

Introduction to geology, mineralogy and sedimentology

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2. Development of the geological world view
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9. Exogenic processes
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12. Geology of Chile
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3. Literature

Literature:

Gill

Press, F. & Siever, R. (2003): Allgemeine Geologie. Spektrum Akadem. Verlag. Heidelberg.
3. Auflage, 723 S., ISBN: 3827403073.

Rollinson

Skinner, B.Y. and Porter S.C. (2004): The Dynamik of the Earth: An Introduction to physical geology. 5. Auflage, 648 Seiten. ISBN: 0-471-15228-5.

Tucker, M. (1996): Methods of Sedimentology. 366 S. Enke-Verlag.

Web sites related to mineraly, rocks and geology:

<http://www.seilnacht.com/Minerale/index.htm> (Minerals)

<http://www.min.uni-bremen.de/kabinett/> (Minerals)

<http://www.geolab.unc.edu/Petunia/lqMetAtlas/mainmenu.html> (rocks)

<http://www-seismo.hannover.bgr.de> (Seismik weltweit)

<http://pubs.usgs.gov/gip/dynamic/dynamic.html> (plate tectonics)

<http://www.wikipedia.de>

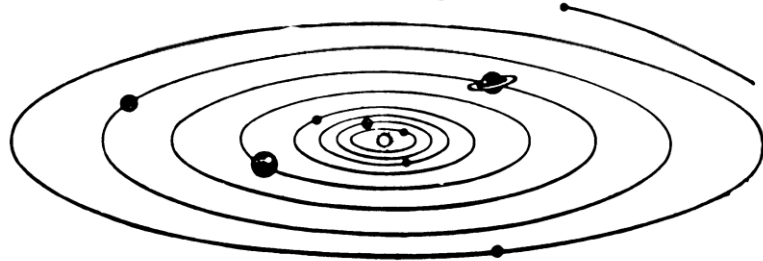
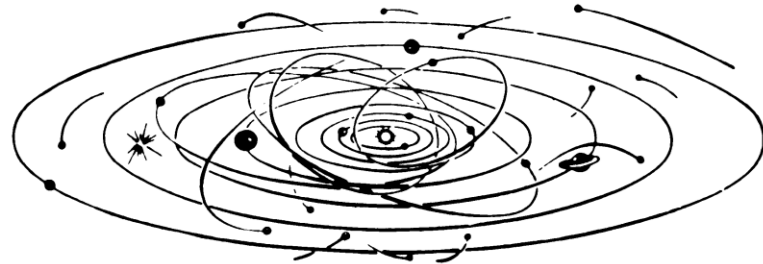
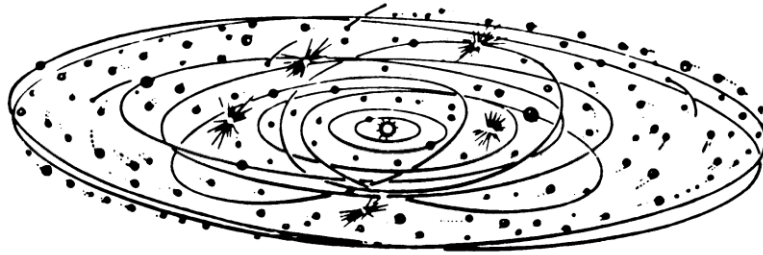
4. Formation our solar system from a planetary cloud of dust (here Orion Nebula)



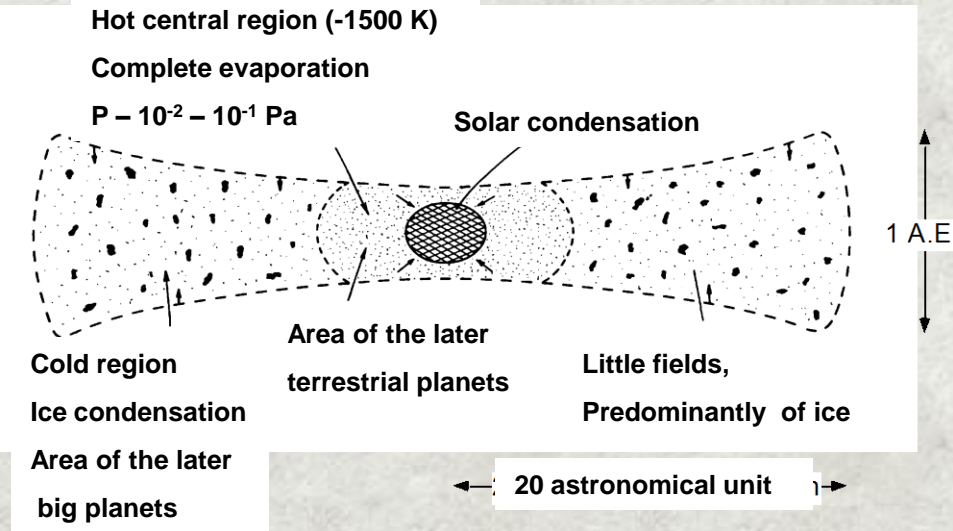
Early history of the inner Planets and the origin their Schalenbaus

- The solar system was formed from by condensation of a dust nube (first homogeneous later fractionated condensation → 10 GA)
 - 99% of the mass in the sun only 1% in satellites.
- *Planetary simals* and the later *Protoplanets* moved to Keplerbahnen orbits
 - 100 Ma gravitative cleaning of the orbits around the sun

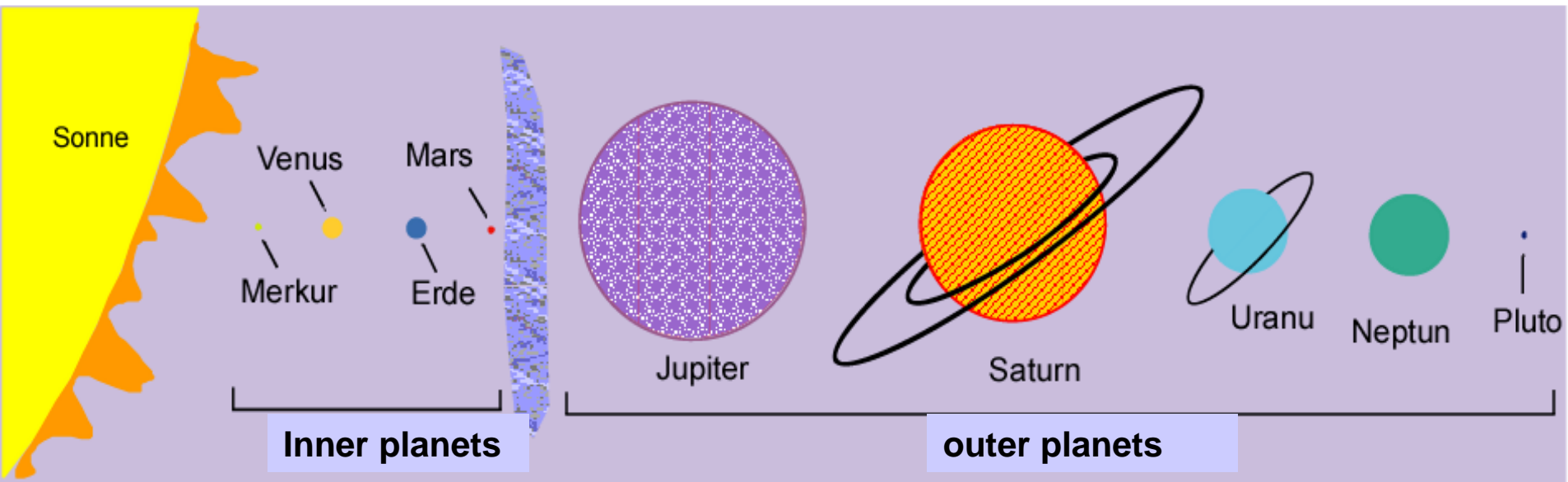
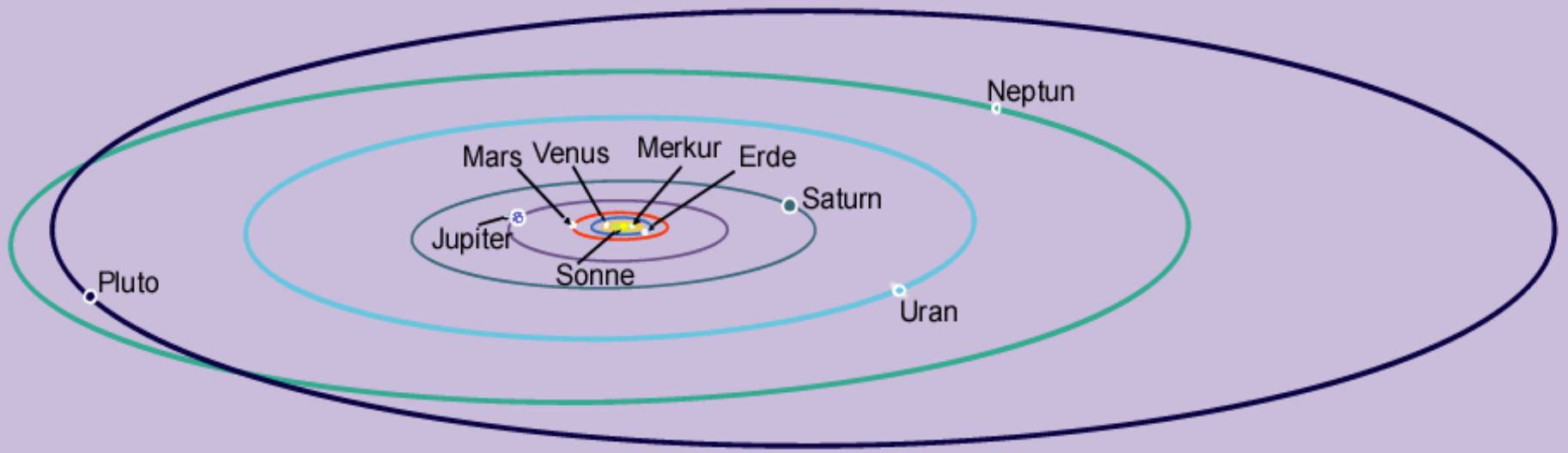
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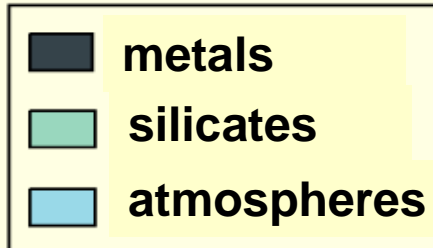


from unpublished Ms, Rainer Altherr

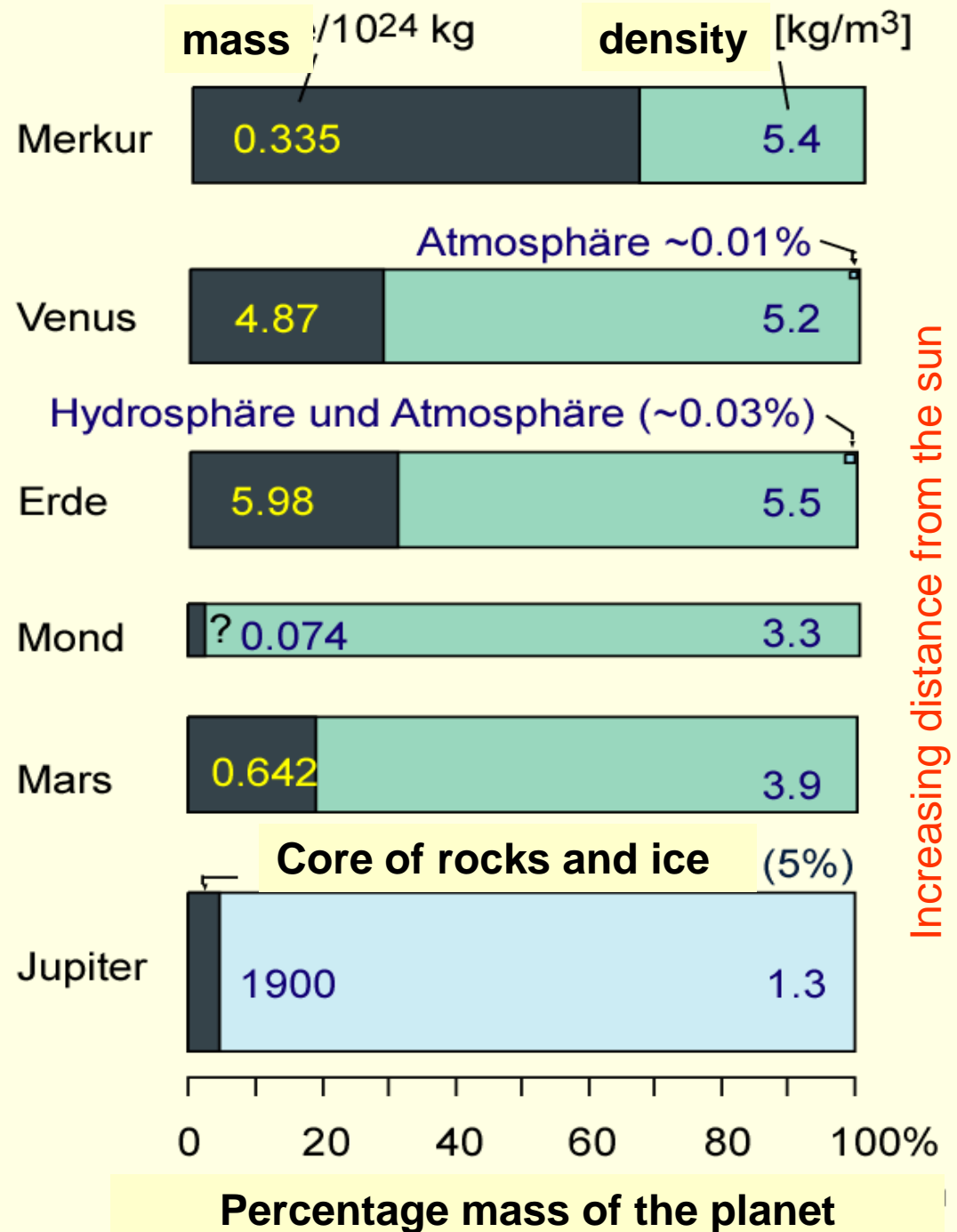


Distances from the sun, relative diameters of planets

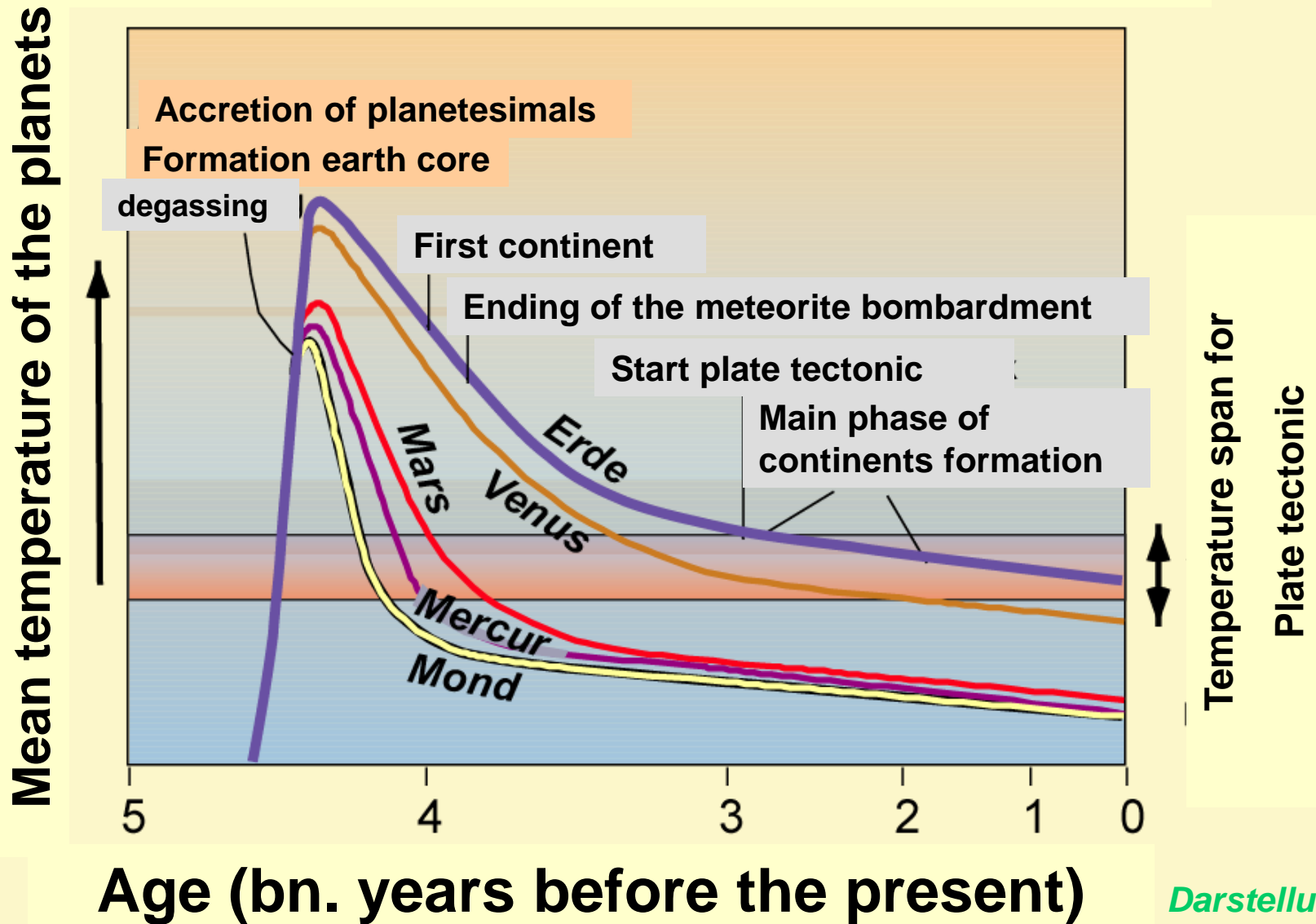
Composition of planets of our solar system



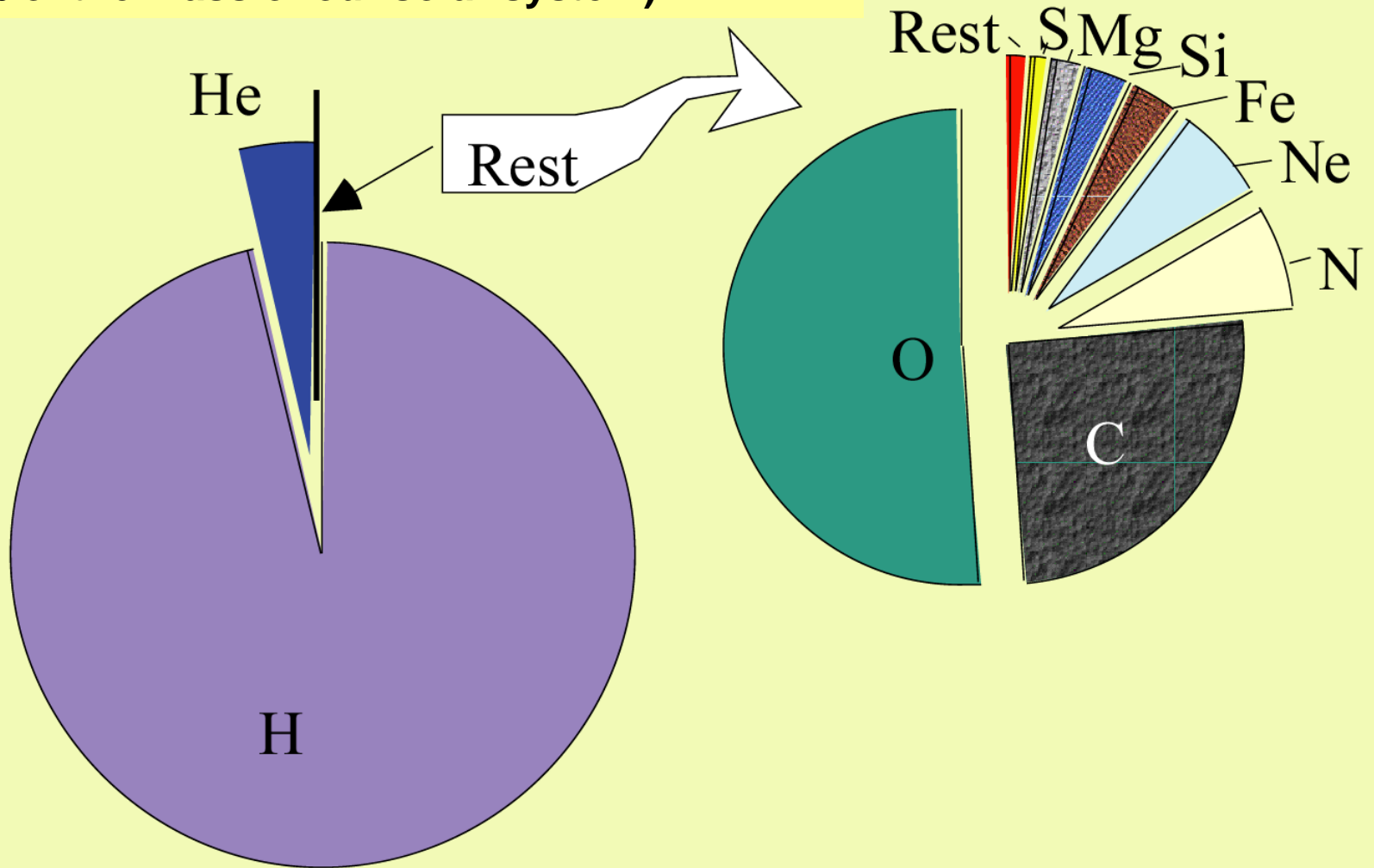
Gill (1994)



Temperature development of the planets since their formation



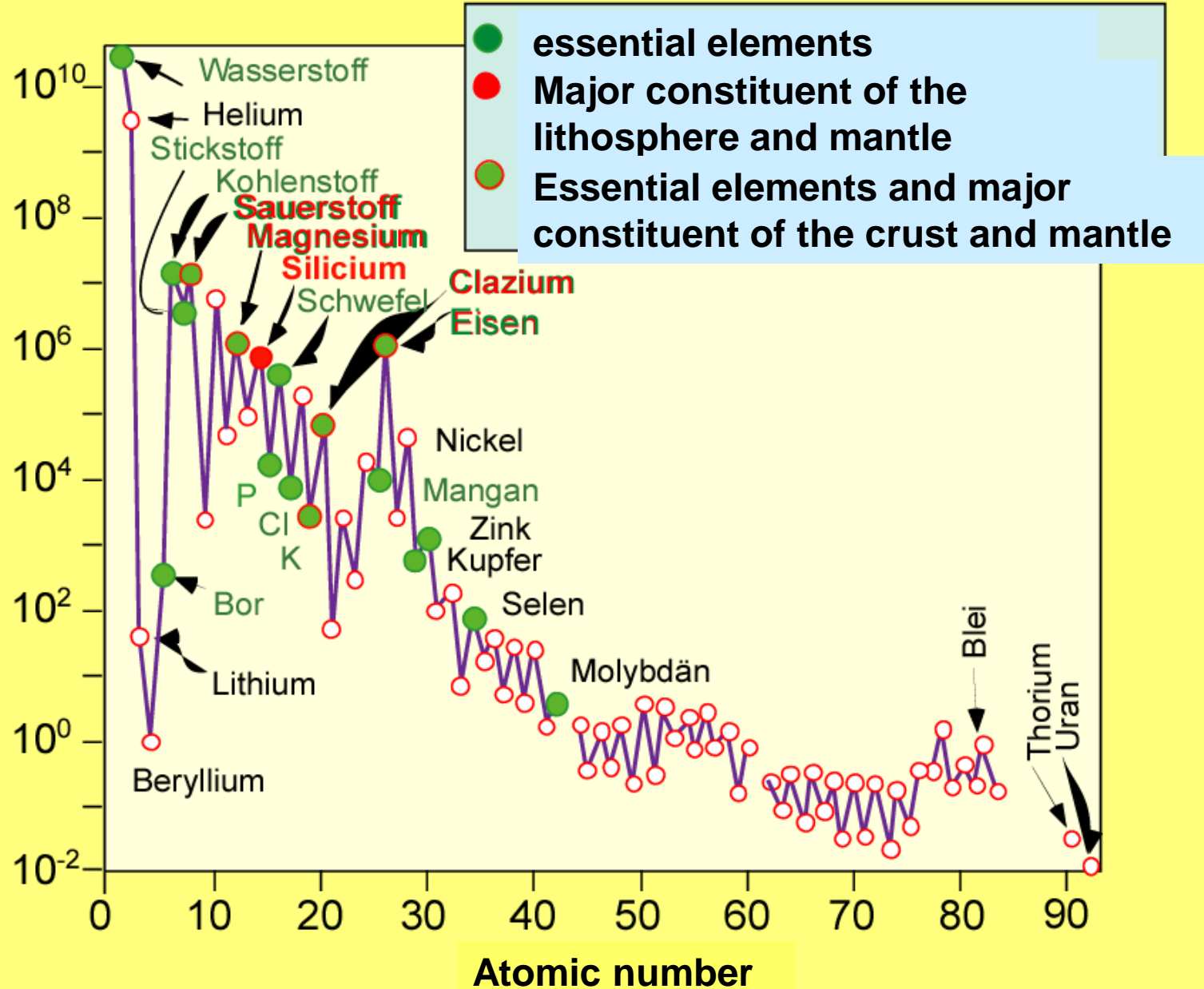
**Chemical composition of the sun
(=98% of the mass of our solar system)**



Element occurrence in the Solar system

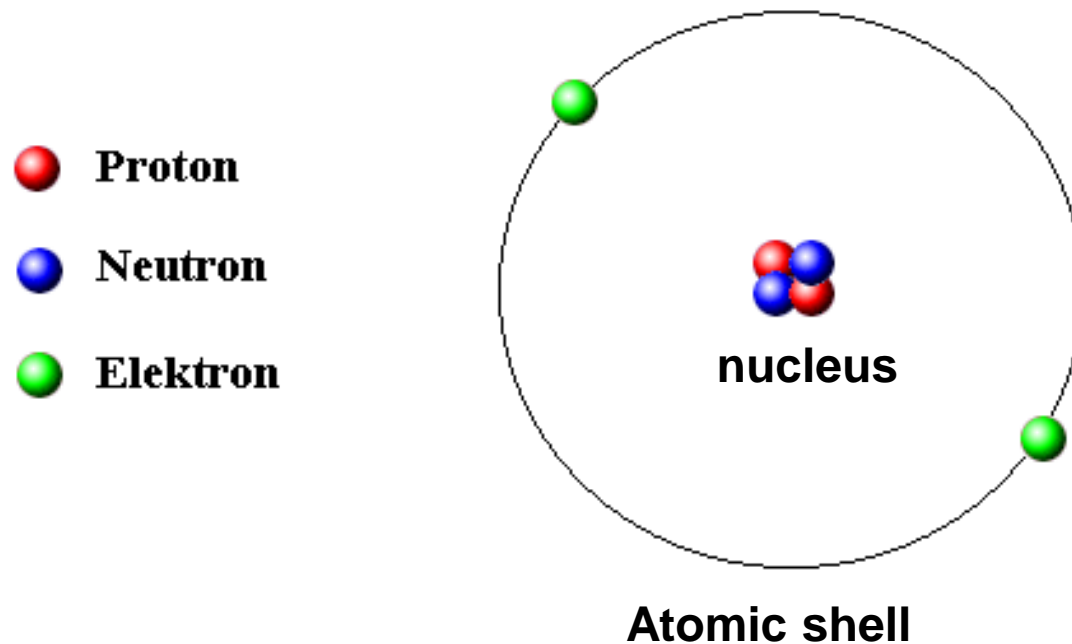
Occurrence of the element in the solar system

relative to 10^6 Si-atoms



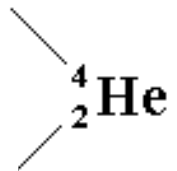
Atomic Structure:

Heliumatom: ${}^4_2\text{He}$



Nucleon number - mass number
(protons and neutrons together)

Atomic number
(number of the protons)



Gruppe I	Gruppe II	Significant elements for the mineralogy										Gruppe III	Gruppe IV	Gruppe V	Gruppe VI	Gruppe VII	Gruppe VIII
H 1,01 +1 Wasserstoff																	He 4,0 Helium
Li 6,94 +1 Lithium	Be 9,01 +2 Beryllium											B 10,81 +3 Bor	C 12,01 -4 +4 Kohlenstoff	N 14,01 -3 +5 +3 Stickstoff	O 16,0 -2 Sauerstoff	F 19,0 -1 Fluor	Ne 20,18 Neon
Na 22,99 +1 Natrium	Mg 24,31 +2 Magnesium											Al 26,98 +3 Aluminium	Si 28,09 +4 Silicium	P 30,97 -3 +3 Phosphor	S 32,07 -2 Schwefel	Cl 35,46 -1 Chlor	Ar 39,95 Argon
K 39,10 +1 Kalium	Ca 40,08 +2 Calcium	Sc 44,96 +3 Scandium	Ti 47,88 +3 +4 Titan	V 50,94 +2 +3 +4 +5 Vanadium	Cr 52,0 +2 +3 +4 Chrom	Mn 54,94 +2 +3 +4 Mangan	Fe 55,85 +2 +3 Eisen			Ni 58,69 +2 +3 Nickel	Cu 63,55 +1 +2 Kupfer	Zn 65,39 +2 Zink			As 74,92 +3 +5 Arsen	Br 79,9 -1 Brom	Kr 83,8 Krypton
Rb 85,47 +1 Rubidium	Sr 87,62 +2 Strontium	Y 88,91 +3 Yttrium	Zr 91,22 +2 +4 Zirkonium	Nb 92,91 +2 +4 +5 Niob	Mo 95,94 +2 +6 Molybdän						Ag 107,87 +1 Silber			Sb 121,75 +3 +5 Antimon		I 126,9 -1 Jod	Xe 131,29 Xenon
Cs 132,91 +1 Caesium	Ba 137,33 +2 Barium		Hf 178,49 +4 Hafnium	Ta 180,95 +3 +4 +5 Tantal						Pt 195,08 +2 +4 Platin	Au 196,97 Gold	Hg 200,59 +1 +2 Quecksilber		Pb 207,2 +2 +4 Blei	Bi 208,98 +3 +5 Bismut		Rn 222,02 Radon

Lanthanoide
(Seltene Erden)

La 138,91 +3 Lanthan	Ce 140,12 +3 +4 Cer		Nd 144,24 +3 Neodym		Sm 150,36 +3 Samarium	Eu 151,97 +2 +3 Europium									Yb +3 Ytterbium
	Th 132,04 +3 +4 Thorium		U 238,03 +3 +4 +5 +6 Uran												

Actenoide

Alkalimetalle	Erd-Alkalimetalle	Ü b e r g a n g s m e t a l l e										Metalle	Halbmetalle	Nichtmetalle	Halogene	Edelgase
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 wichtige Elemente der Erdkruste

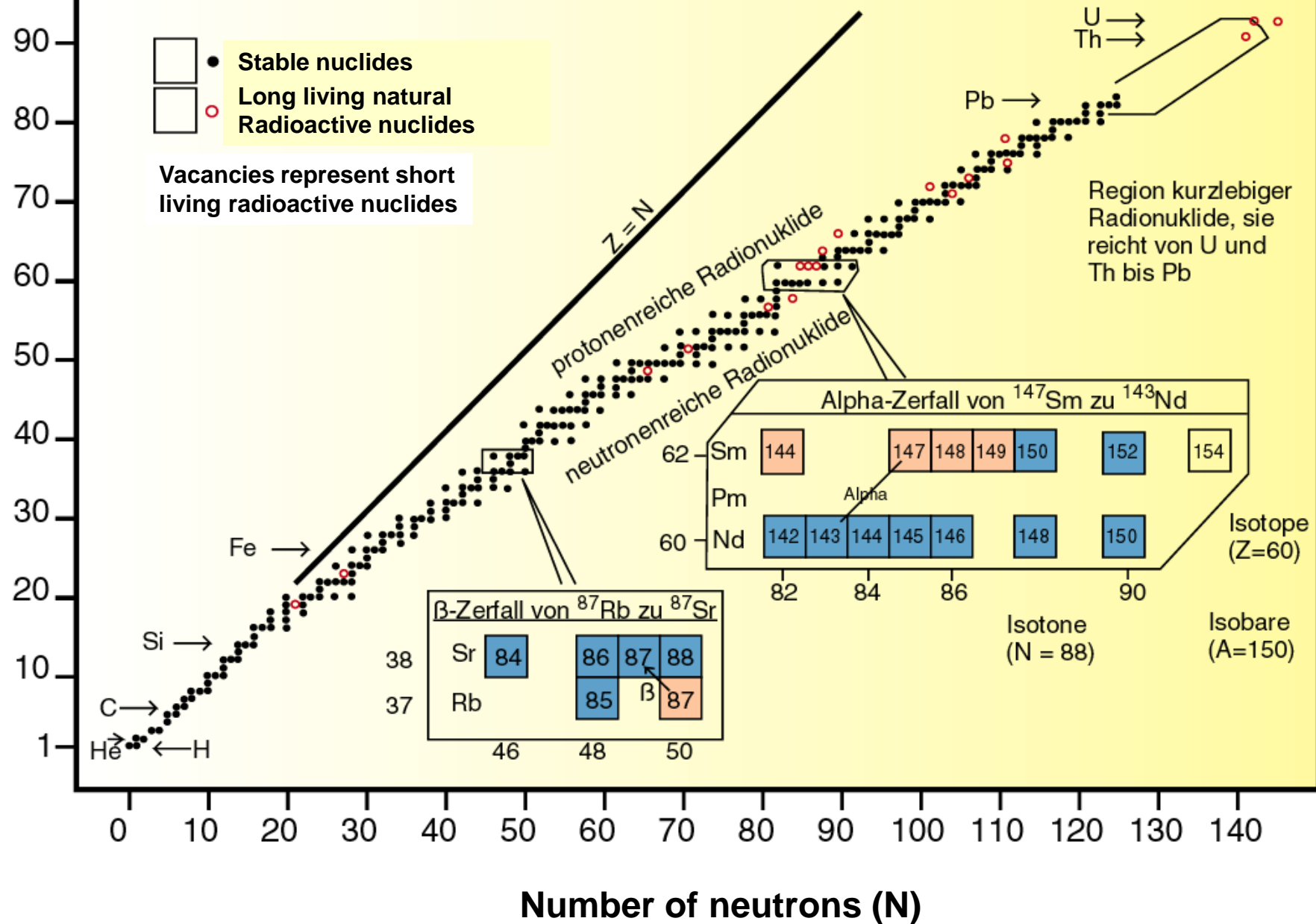
 radioaktiv

 gediegen

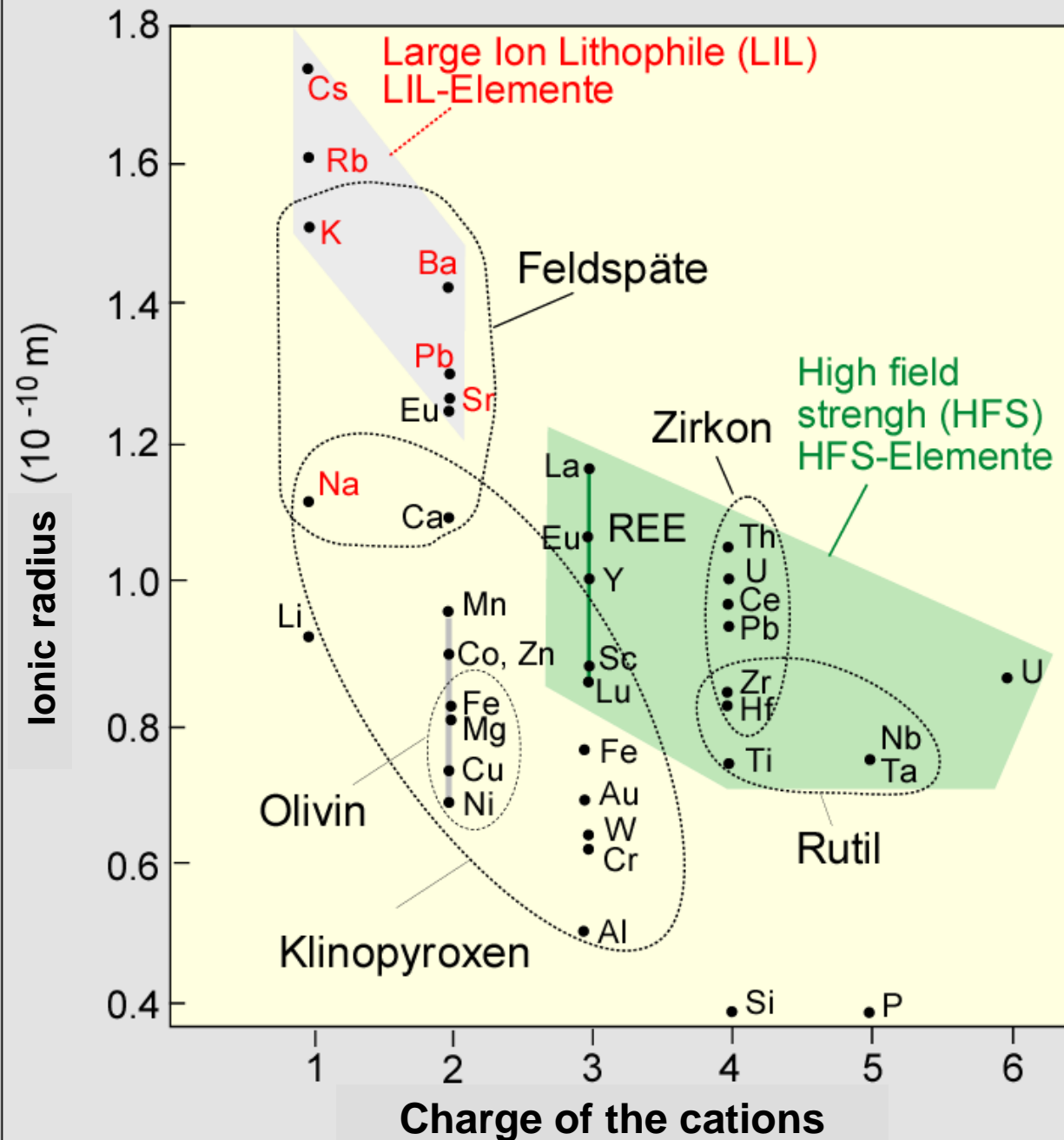
85,42 durchschnittliche relative Atommasse

14 radioaktives Isotop

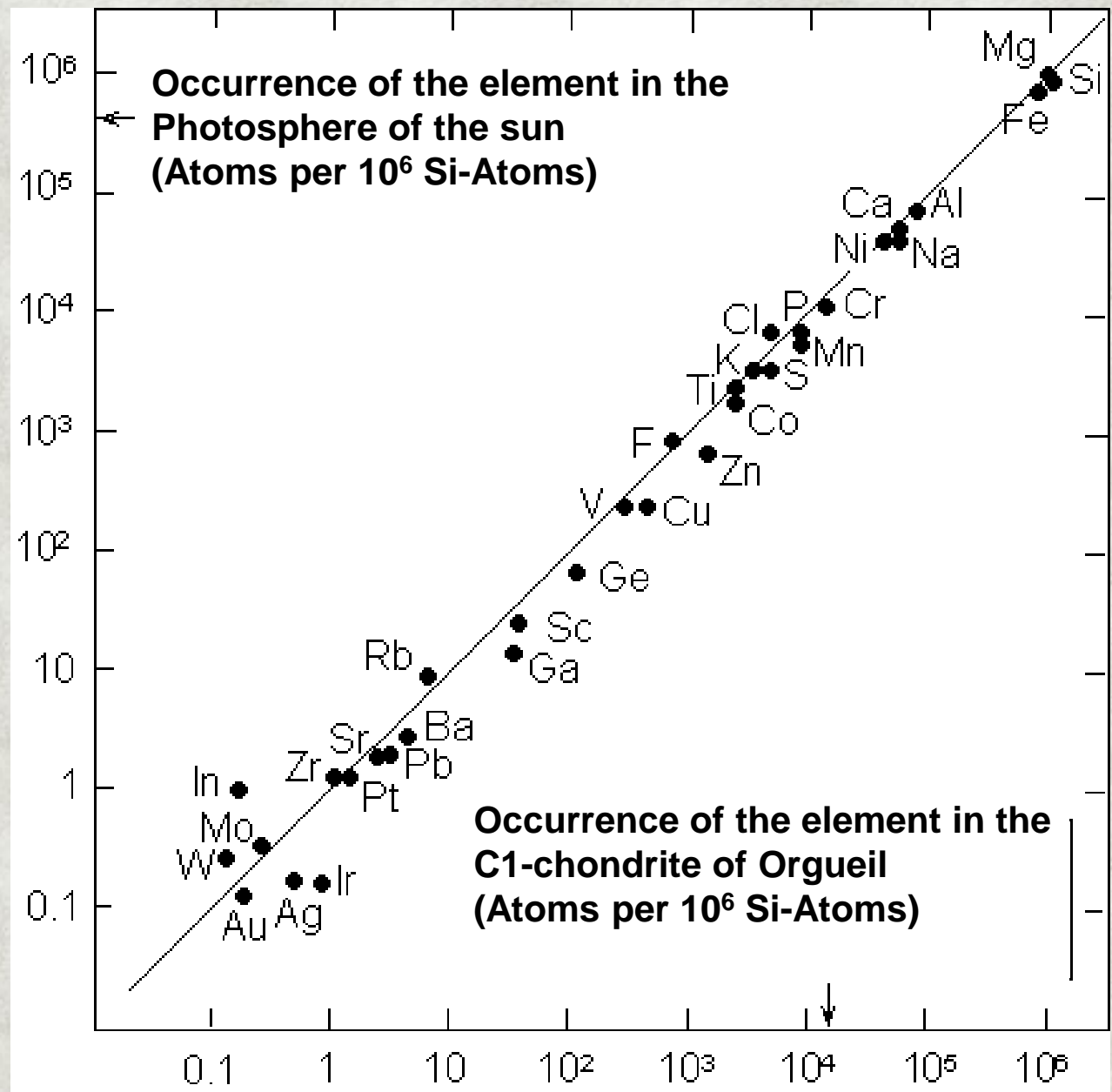
Number of protons (Z)



The minerals olivine, zircon, feldspars, clinopyroxene and rutile have different tolerance (pointed curve) relative to the ionic radius and the charge of the cations, by possible assembly of this cations in the crystal lattice of a mineral



Comparison of
the **element**
concentrations in
the photosphere
of the sun
(Spectral lines of
single elements)
with that of the
C1-chondrite of
Orgueil.



6. Meteorites

Definition: Meteorite

natural object, which comes from the outer space and enters in the atmosphere or in the gravity field of the earth or other celestial bodies (in particular a planet or a moon). The meteorite can there to burn up or to reach the surface. (→ Crater formation, more than 200 are known)

The meteorites hail from the solar system and are well remains of asteroids (planetoids or minor planets).

Meteor,

the term meteor comes from the Grecian word *meteoron*, which mean “high in the air”. The small particles of matter in the solar system that are directly observable only by their incandescence from frictional heating on entry into the atmosphere.

Meteorites

- *Undifferentiated meteorites* from millimeter size spherical silicate aggregate, the *chondres*, → *Chondrite*
→ document the early composition of the solar system
- *Differentiated meteorites* of planets, which are differentiated by dissolution in core, mantel and crust. In this group are the *Achondrite* (Stone-Meteorite), *Stone-Iron-Meteorites* and *Iron-Meteorite*.
 - Document the evolution of the inner planets of the solar system.

Meteoriten Typen

Iron: to be composed basically of iron and nickel;
similar to M Asteroid;



Stone-Iron: to be composed of iron and Stone-
material; similar to S Asteroid;



Chondrite: by far the more abundant; the composition
is similar to the mantle and crust from planets.



Meteoriten Typen

Carbonaceous Chondrite: similar to C Asteroidean;



Achondrite: similar to basalt stone; origin → the Moon and Mars



Meteorite statistic		between 1740 and 1990 (total 4660piece); only 33% are KAT (1)				
Typ	observed	[1] %	casual encountered	[2] %	[1] weight	[2] weight
Stone	95.0		79.8		15200 kg	8300 kg
Stone-Iron	1.0		1.6		525 kg	8600 kg
Iron	4.0		18.6		27000 kg	435000 kg

Meteorites



Achondrite-Eukrit (Stone meteorite)
Camel Donga Museum (Australia)



Achondrite Camel Donga Museum (Australia)

Picture from http://www.carionmineraux.com/meteorite_chondrites.htm

Meteoriten

C1 Chondrite Allende

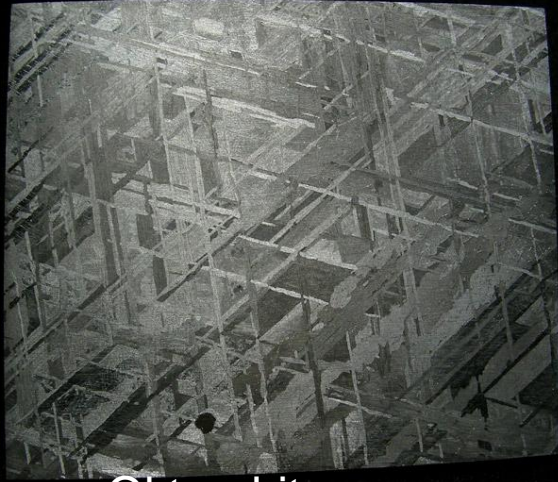


Undifferentiated Meteorite

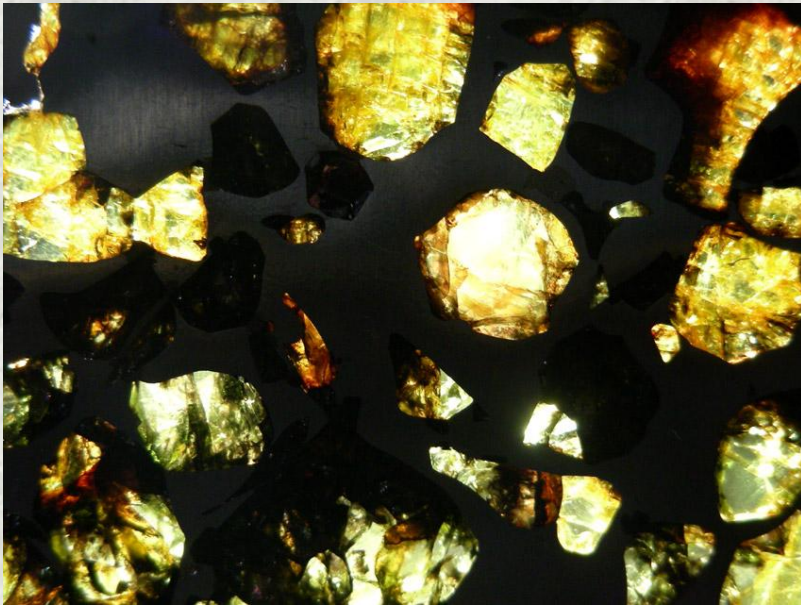


Achondrite Tatahouine (Stone meteorite)

Iron meteorite and Stone-Iron meteorite



Oktaedrite



Iron meteorite with Olivine (Esquel)

Iron meteorite: Core destroyed planets





Ca-Al-inclusions in carbonate Chondriten:
 4567 ± 0.6 Million Years alt

protosolarer
Nebel

undifferenzierter
Asteroid

differenzierter
Asteroid

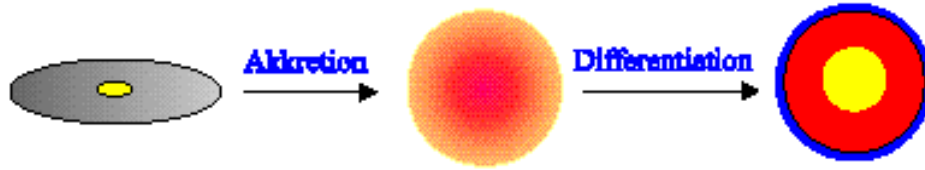


Abb. 1.7: Ursprung von undifferenzierten und differenzierten Meteoriten

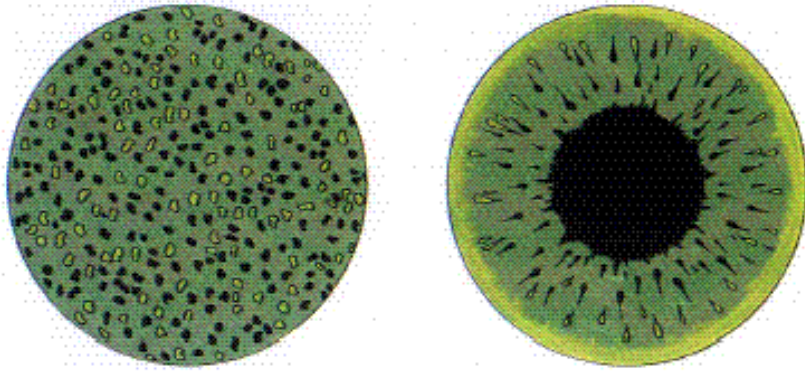
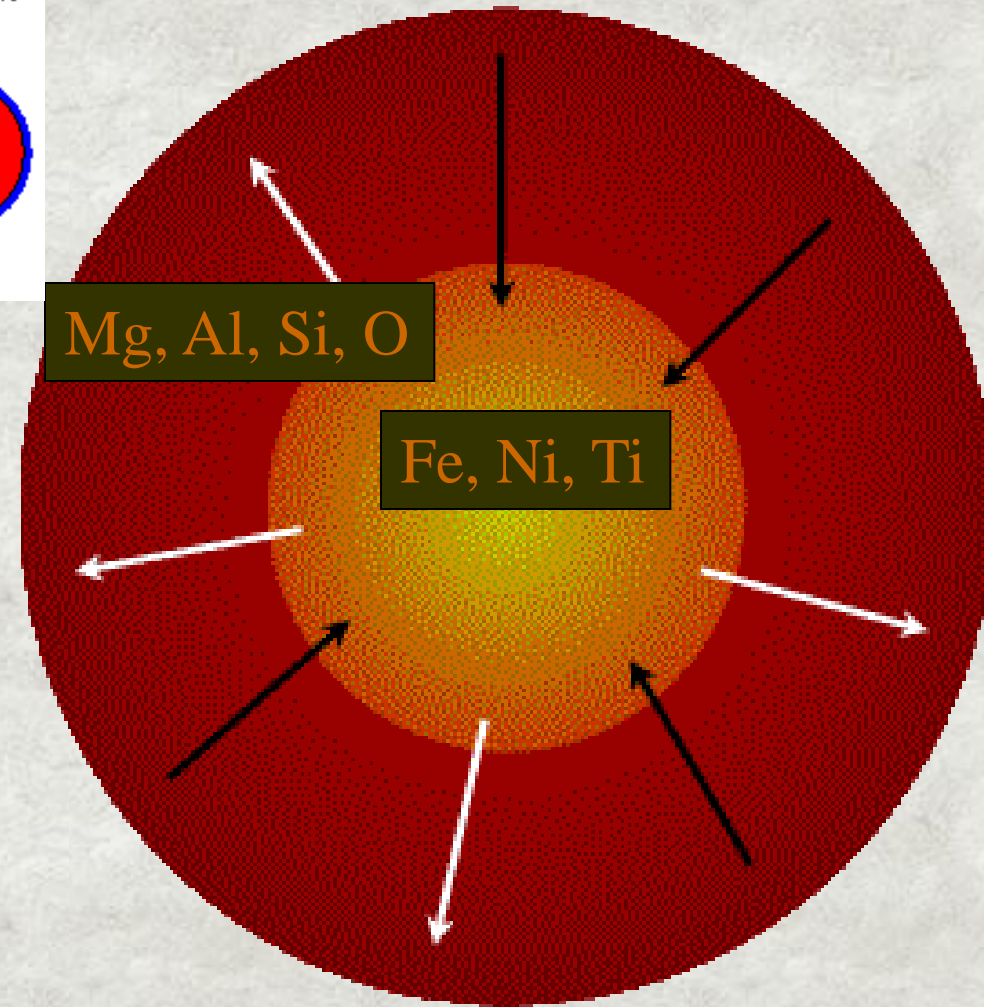


Abb. 1.16: Differentiation eines Asteroiden. Links: undifferenzierter Körper mit Metall u. Sulfid (schwarz) und niedrigschmelzenden, alkalireichen Silikaten (gelb), sowie hochschmelzenden Mg-Silikaten (grün). Rechts: Bildung von Kern und Kruste - die schweren metallischen und sulfidischen Komponenten sinken nach unten, die leichten, silikatischen Komponenten steigen auf und bilden die Kruste.



Differentiation some Asteroids:

and all Planets: heavy elements as Iron and Nickel sink in the Core (black arrows), light elements as Magnesium, Aluminum und Oxygen ascend and build the Mantel (white arrows).

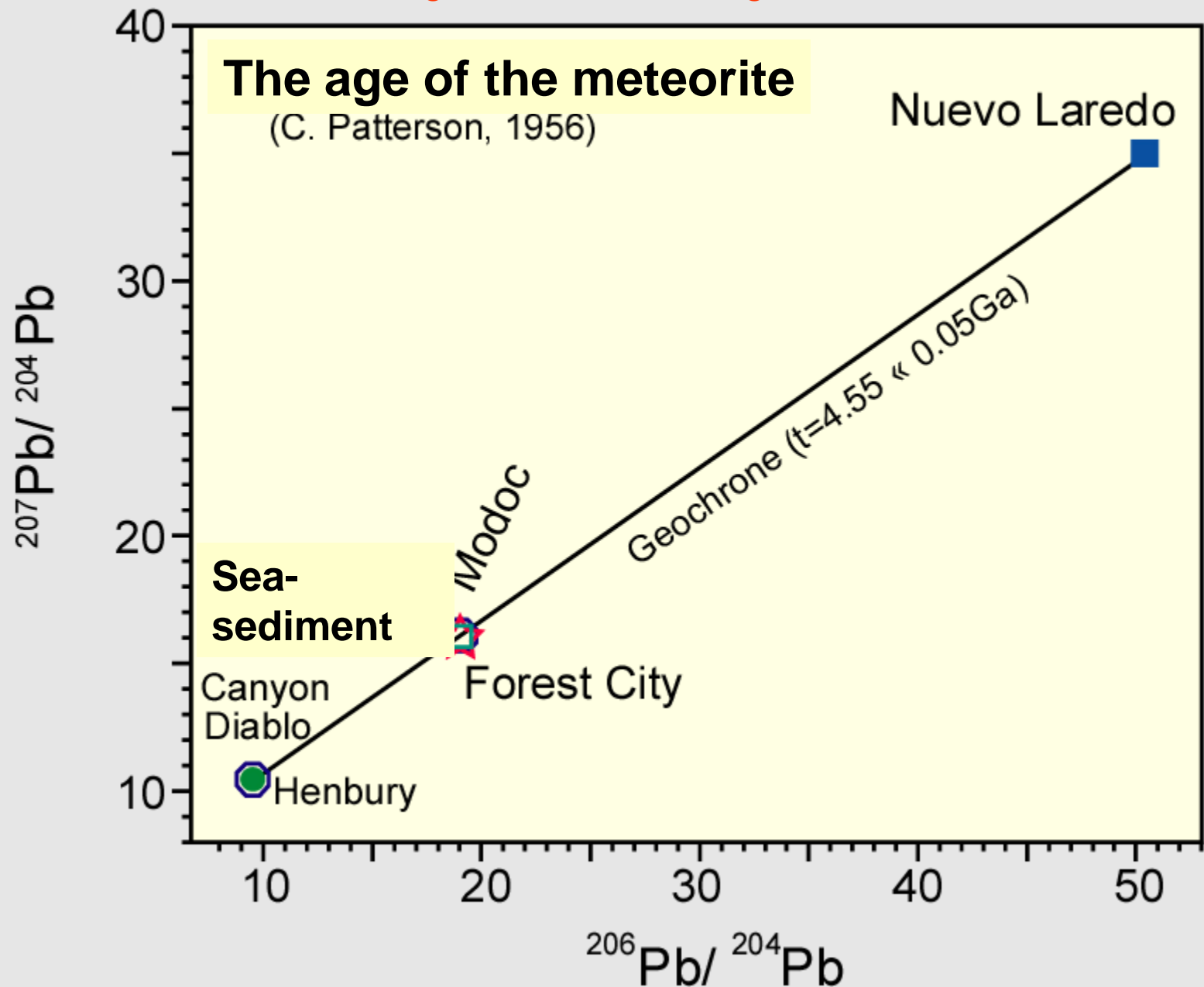
„Meteorites capture“

At the beginning of the earth's history from 4.6 Ma impacted more of hundred millions years long over the Earth, Moon and other planets numerous meteorits and asteroids.

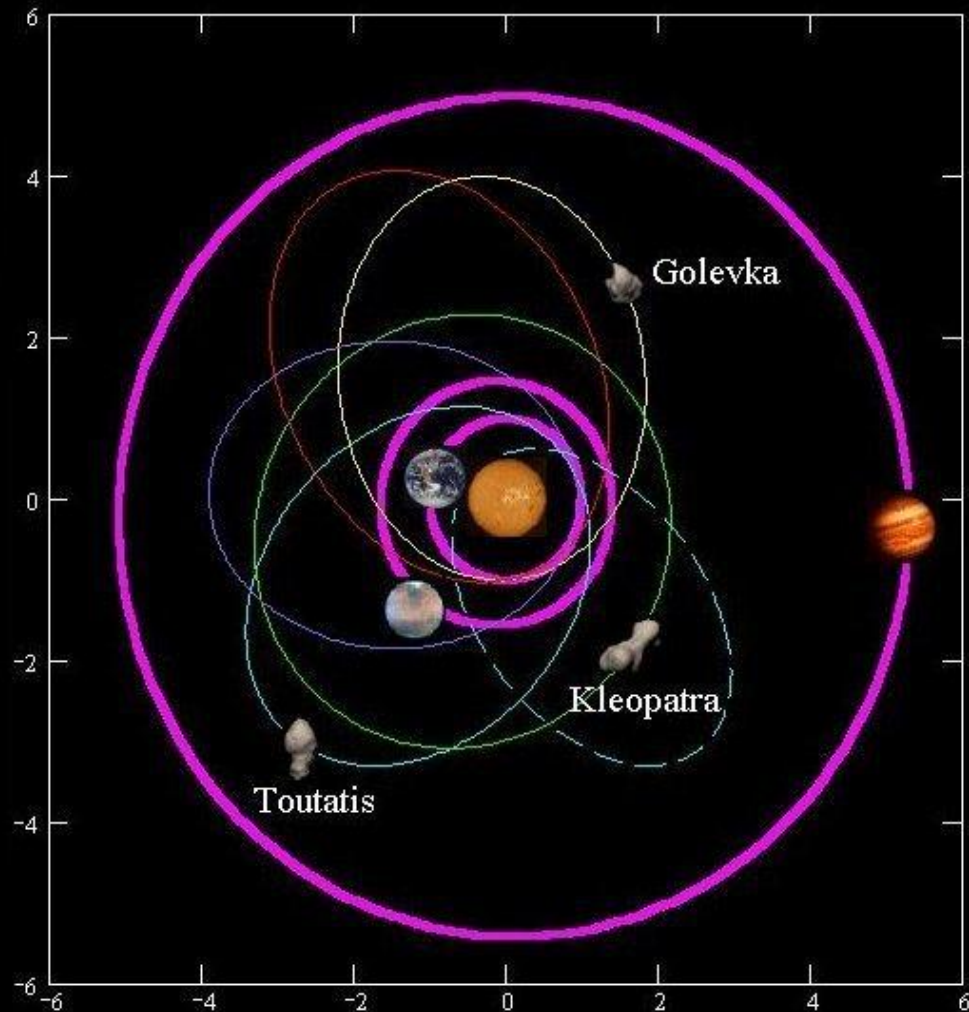
Due to tectonic movements and the posterior stark alteration and erosion haven not conserved vestiges for this meteorite's hail on the earth, whereas the Moon and Mercury show still many craters.

From 3.8 Milliard years declines the meteorite's hail.

Something different for the age of the meteorite



Earth Crossing Asteroids Main Menu

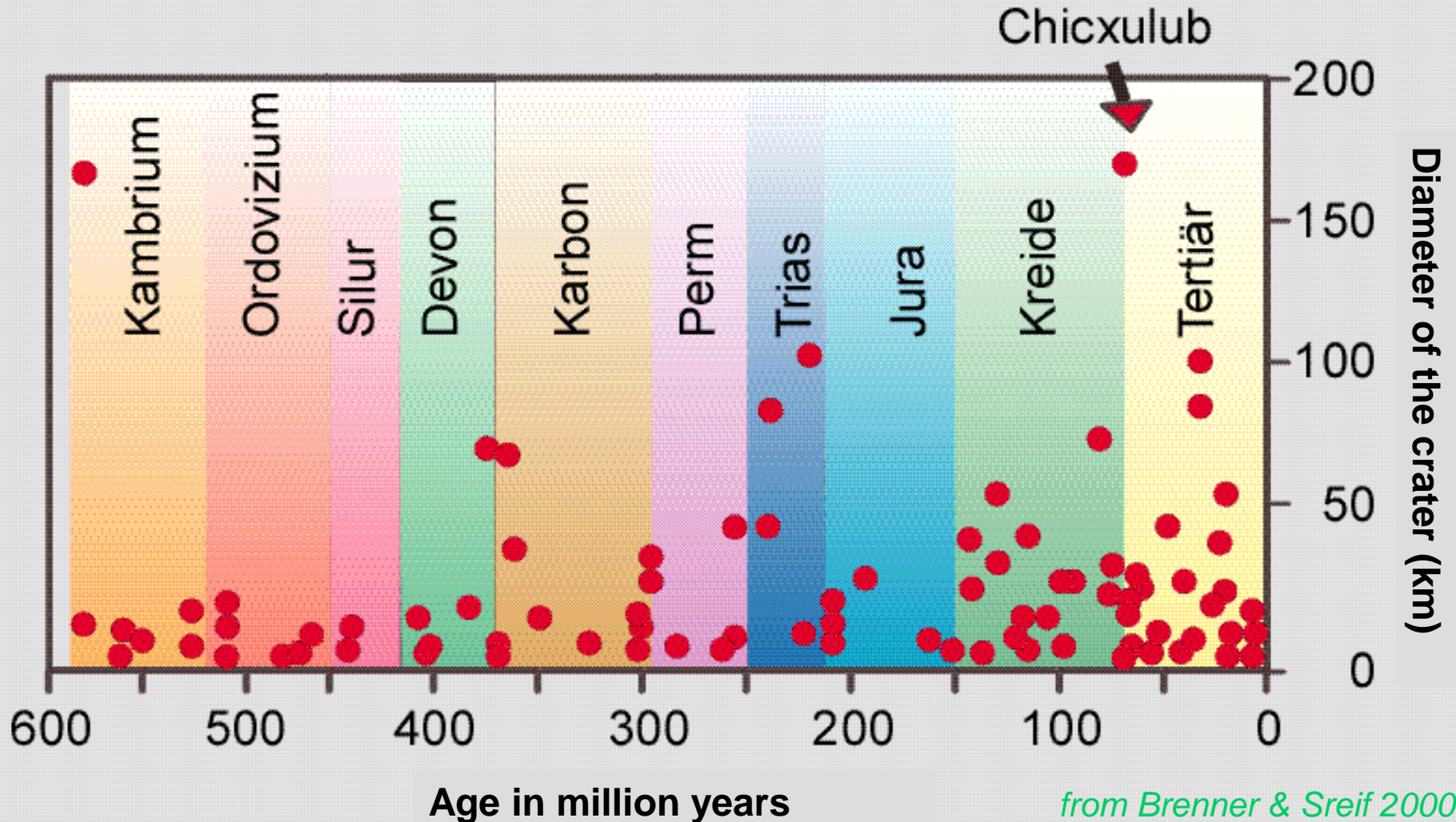


ersichtliche
Asteroidenbahnen welche
die Erdbahn kreuzen



Distribution at present still existing structures, which can explained by impact of meteorites.

Relationship between Earths epochs and meteorite impacts



Large Meteorite impacts:

590 Million years ago: The **Acraman crater in South Australia** with 160 km diameter documents a large impact

367 Million years ago the **Alamo Impact Breccia** in Nevada USA, documents a strong impact.

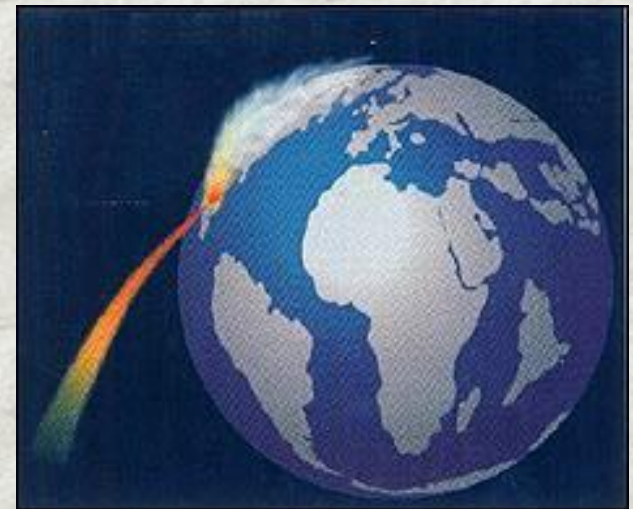
220 Million years ago the 80 km wide **Puchez-Katunki crater** in Russia, NW of Nishnij Nowgorod, was produced

Yukatan bzw. Chicxulub Meteorite

65 Million years ago a strong impact occurred at Peninsula Yucatán, in Chicxulub, in Mexico. It marks the extinction of many of the dinosaurian.

The cretaceous-Tertiary- transition is characterized by the occurrence of the metal Iridium (positive Iridium anomaly).

The disturbances on the environment by this meteorite impact is still strongly disputed.



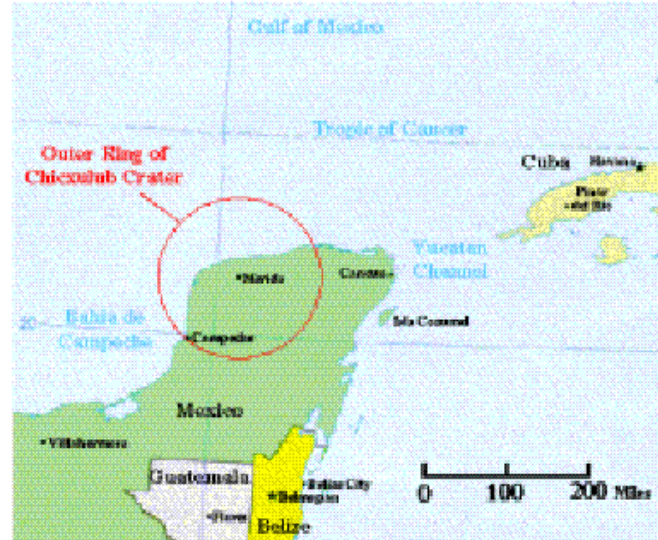
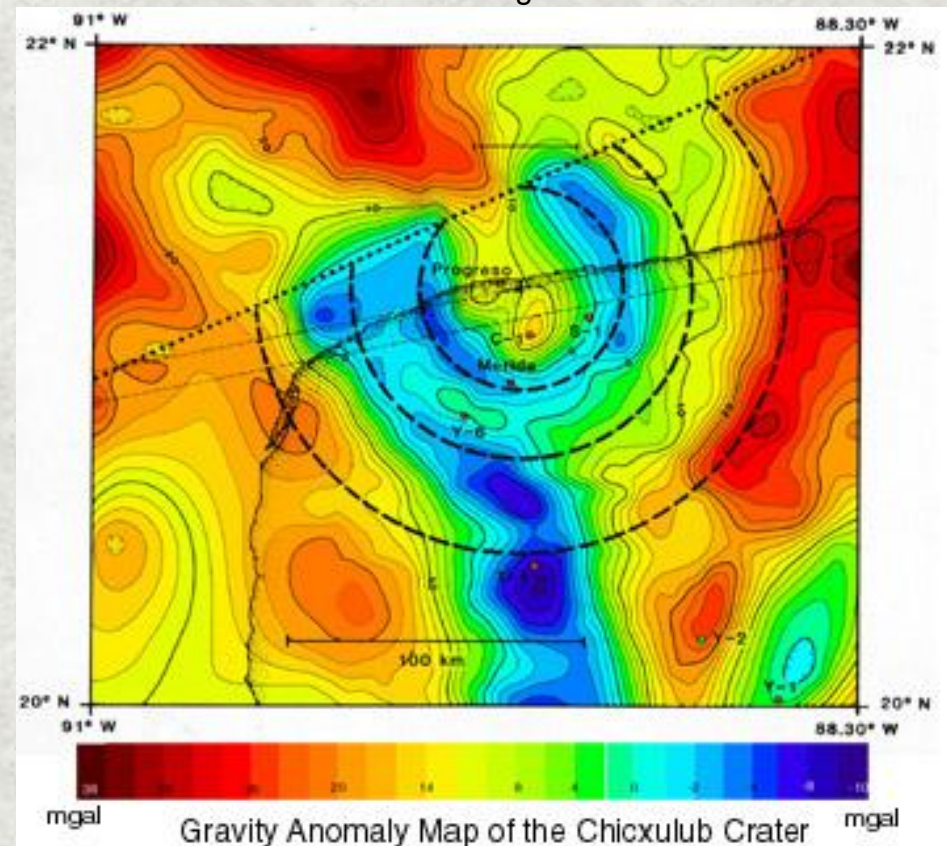
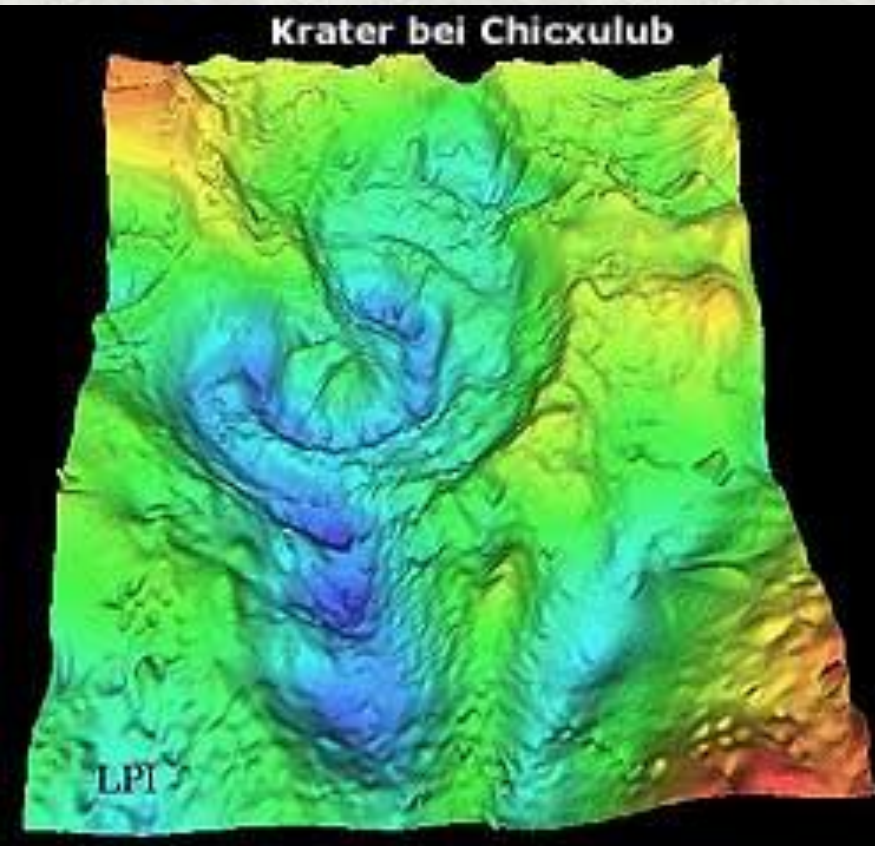


Abb. 1: Das Einschlagsgebiet auf der Halbinsel Yukatan (Mexiko)¹

The impact structure seen as gravity anomaly

In view of the tremendous energies involved, it is no wonder then that we classify the Chicxulub impact in the Yucatan Peninsula as one of the biggest short-term natural events known in the geologic record (of nuclear-equivalent magnitude in excess of 100 trillion tons of TNT equivalent). It occurred 65 million years ago and led to a 200-300 km (>150 mi) wide (there's still some uncertainty regarding the location of the outer rim) and perhaps 16 km (10 mi) deep depression.

This huge structure has no evident surface expression, being covered by younger sedimentary rocks, but does appear subsurface as a strong gravity anomaly, as shown below. It was discovered almost incidentally through oil drilling, in which core samples, containing so-called volcanic rocks (now known to be shock-melted rock), showed distinct shock effects. The samples languished for years in the basement of the University of New Orleans' Geology Building, before someone re-examined them and discovered their origin.



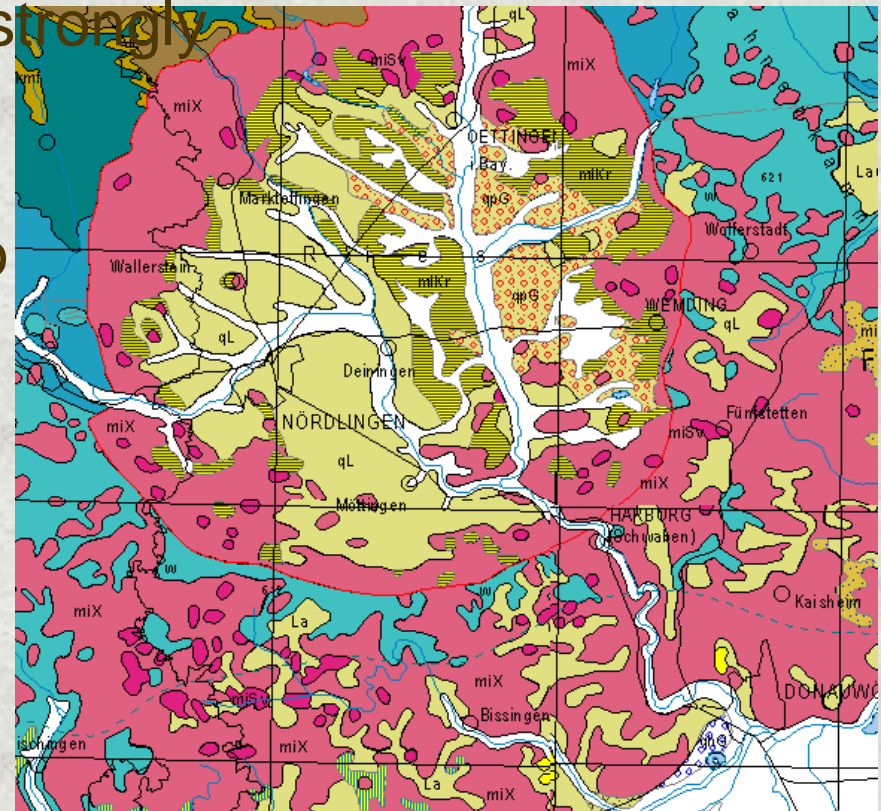
Demise of the dinosaurs

The Chicxulub Crater is believed to be the result of the collision with an asteroid measuring some 10 to 20 km across. The environmental effects that accompanied its formation were thought to have been responsible for the mass extinction at the end of the Cretaceous period, about 65 million years ago, in which the last of the dinosaurs, along with many other species, disappeared (see Cretaceous-Tertiary Boundary). However, this theory was called into question in February 2004, when An international group of scientists led by Professor Gerta Keller, of Princeton University, published results, based on a core sample, that the Chicxulub crater predates the extinction of the dinosaurs by about 300,000 years. The authors say this impact did not wipe out the creatures, rather two or more collisions could have been responsible. Keller and her colleagues analyzed rock from their core using five separate indicators of age, including fossil planktonic organisms and patterns of reversals in the Earth's magnetic field. The results suggest the crater was punched into the Earth around a third of a million years before the dinosaurs disappeared from the face of the planet. Keller and her team contend their findings prove the Chicxulub impact did not by itself trigger the extinction of the great beasts. Instead, they believe a cooling of the global climate shortly followed by a period of greenhouse warming placed enormous stress on the dinosaurs. This warming could have been triggered by carbon dioxide released by a massive eruption of lava seen today in the Deccan traps of India. The Chicxulub impact occurred during this warming period and, although the environmental effects were severe, it did not cause the extinction of the dinosaurs. The team believes a second impact, 300,000 years after the Chicxulub collision, finished off the creatures. The structure of the sea bed beneath the Indian Ocean suggests this second impact could have been there, Keller has indicated.

Further details see: http://en.wikipedia.org/wiki/Chicxulub_crater

Impact example of the Nördlinger Ries in Germany

- 15 Million years ago
- 25 km diameter of the crater (see geology map below)
- The discovery of the high pressure quartz modification coesite and Stishovite confirmed that there was an impact. This was disputed strongly previously, like in many other cases where the old impact structure was not so evident.



Geophysical anomalies Anomalien

Gravity anomaly



Popigai crater, Siberia.

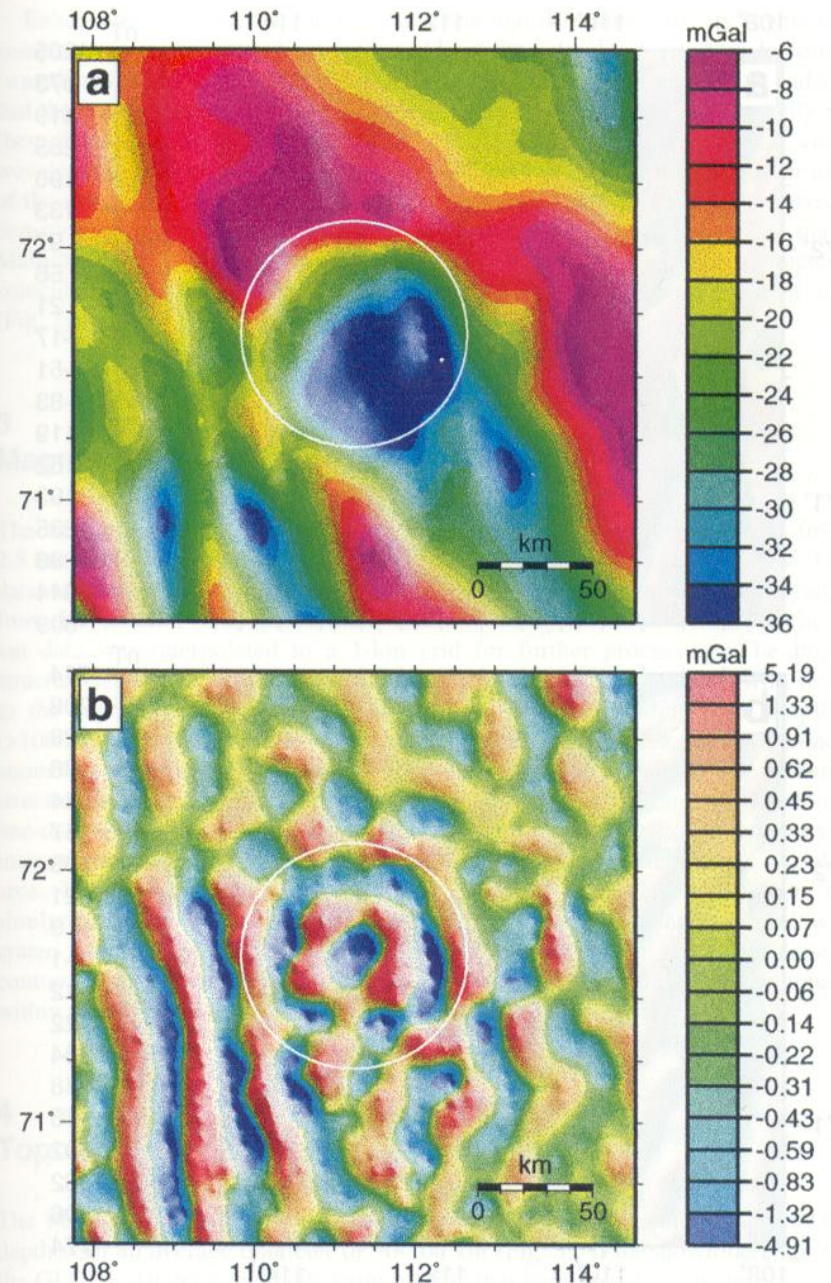


Fig. 3. (a) Gravity field over the Popigai region (data courtesy of GETECH Ltd.). White circle represents estimated crater rim at 100 km diameter. (b) Gravity field data, high-pass filtered to retain wavelength components <50 km (data courtesy of GETECH Ltd.).

The Tunguska meteorite impact

- Almost a century ago, on June 17 (30), 1908, a massive explosion occurred near the Podkamennaya Tunguska River, in what is now Russia's Krasnoyarsk Territory, Central Siberia. The residents of the Vanavara trading post, 65 kilometers (40 miles) south of the blast site, later claimed that the ground trembled violently when attacked by a huge ball of fire, followed by a terrible storm that destroyed everything in its wake.
- The explosion was most likely caused by the airburst of a large meteoroid or comet fragment at an altitude of 5-10 kilometers (3-6 miles) from the Earth's surface
- Studies have yielded varying estimates for the object's size, with most experts agreeing that it measured several dozen meters in diameter. Estimates of the energy of the blast range from 5 megatons to as high as 30 megatons of TNT, with 10-15 megatons being the most likely yield. The blast, about 1,000 times more powerful than the Hiroshima bomb, felled an estimated 80 million trees over 2,150 square kilometers (830 square miles). The earthquake caused by the blast measured 5.0 on the Richter scale. The region has never completely recovered.
- The Tunguska blast was the largest meteoroid impact in the Earth's recent history, and demonstrated the awesome destructive power of near-space objects. An explosion of the scale of the one in Tunguska could destroy large metropolitan areas. It is this possibility that has helped to spark discussion of asteroid deflection strategies.
- Due to the rotation of Earth, if the collision had occurred 4 hours 47 minutes later, it would have completely destroyed the Imperial Russian capital, St. Petersburg. A little later still, and the Tunguska meteorite would have wreaked chaos and destruction in densely populated Europe.



Wolfe Creek, Australia

19°18'S, 127°46'E; 0.9 km

Diameter of the crater;
age 300 000 years

Wolfe Creek is a well preserved crater in the northwestern dessert of Australia, which is partly overlain by blown sand. The rim is about 25 m above the surrounding plain Rand and the crater floor is depend around 50m. Oxidized remnants of an iron meteorite and glass from an impact fusion have been found.

Photo from *Courtesy of V. L. Sharpton, LPI*



Barringer meteorite crater in Arizona

diameter 1.2 km; Age: 49,000 years

Despite the classical impact morphology, the origin was disputed for a long time until meteorite fragments (Canyon Diablo meteorite) and breccia as well as minerals and rocks affected by a shock metamorphism have been found.

(Courtesy of USGS/D. Roddy and LPI)

